Cooling of Highly Charged lons

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Why do we need to cool HCI between the EBIT and the precision Penning trap ?

• ion temperature for mass measurement:

 $T_i \lesssim 1 \text{ eV/q}$

• EBIT: $T_i \gg 1 \text{ eV/q}$ must be expected

What do we know about ion temperatures in EBITs ?

- REXEBIS: few 10 eV/q
- Oshima et al.: EBIS/ECRIS "generally" > 10 eV/q
- Dresden EBIT: $T_i q = 3 6 \text{ eV/q}$ measured for Ar¹⁶⁺ but low j_e and E_e
- Livermore, evaporative cooling inside EBIT: 10 - 20 eV/q for Dy⁶⁶⁺ (no data ?)
 Penetrante et al.
- Livermore, evaporative cooling & self-cooling during extraction: T_i ≈ 0.1qeV_{trap} ⇒ ≈ 10 50 eV/q
 Marrs, TITAN workshop 2002

R. Marrs, 2002 TITAN workshop

Ion Heating and Cooling

- Ions are heated by Coulomb collisions with beam electrons
 - Heating rate per ion: H_i :

$$H_i = \pi \frac{j_e}{e} \frac{q^2 e^4}{E_e} \frac{2m_e}{M_i} \lambda_{ie}$$

- Example: ¹⁰⁰Sn⁴⁰⁺ in an Intense EBIT, $H_i \approx 5q$ eV/ms (Note: beam space charge potential $\approx 450q$ eV)
- Evaporative ion cooling reduces ion temperature and emittance



- Controlled injection and evaporation of low-Z ions compensates for electron beam heating of high-q ions
- Thermal equilibrium $\Rightarrow T_i \approx 0.1 qeV_{trap}$
- Self cooling during extraction can produce a dramatic reduction in ion temperature

self-cooling requires slow (ms) spills — not suitable in our case

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Evaporative cooling

collisions with beam electrons heat up ion ensemble
light, less tightly trapped ions (e.g. Ne^{10+•}) evaporate removing thermal energy

• heavy, highly charged ions (e.g. Ba⁵³⁺) remain trapped

Ion temperatures from 1000 eV to 30 eV

Doppler width
$$\Delta\lambda/\lambda \approx 1/20.000$$
 (Ba⁵³⁺)

High resolution spectroscopy

AFAIK, not demonstrated yet

Ion temperature on extraction from EBIT

- actual data appears sparse
- no definite conclusions possible
- emittance/temperature measurements from TITAN EBIT will be necessary, and also interesting in general
- For now must assume that HCI have temperatures of 10...100 eV/q

 \Rightarrow additional cooling before precision trap most likely necessary

Techniques for ion cooling

• buffer gas cooling

- well established method for SCI NO (charge exchange)
- resistive cooling
 - well established, fast enough if Q high enough → would require cryogenic operation
 - ion specific tuning of resonant circuit required

electron cooling

- demonstrated for (anti)protons and HCI at $T_i \gtrsim$ few eV/q
- advantage: electrons self-cool via synchrotron radiation
- disadvantage: electron-ion recombination
- positron cooling
 - avoids recombination, but technically more involved (mCi level source)
- ion-ion cooling with light, cool ions (protons, He²⁺)
 - no recombination issues
 - but no synchrotron cooling, need initially cold light ions
 laser cooling?

Electron cooling

• electron self-cool via synchrotron radiation • for $B = 6 \text{ T} \rightarrow \gamma_c \approx 0.1 \text{ s}$

$$\gamma_{\rm c} = \frac{1}{\tau_{\rm e}} = \frac{4e^2\omega_{\rm c}^2}{3m_{\rm e}c^3}$$

$$\omega_{\rm c} = eB/m$$

Spitzer formula for equilibration time:
 need to take into account continuing electron self-cooling

$$\tau_{\rm i} = \frac{3m_{\rm e}m_{\rm i}c^3}{8\sqrt{2\pi}n_{\rm e}Z^2e^4\ln(\Lambda)} \left(\frac{kT_{\rm i}}{m_{\rm i}c^2} + \frac{kT_{\rm e}}{m_{\rm e}c^2}\right)^{\frac{3}{2}} \qquad \ln(\Lambda) = \ln\left(\frac{b_{\rm max}}{b_{\rm min}}\right)$$

J. Bernard et al. | Nuclear Instruments and Methods in Physics Research A 532 (2004) 224–228



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Problem: recombination

• radiative recombination

- $\circ \sigma \propto 1/E_r \rightarrow \text{significant if electron and ion velocities match}$
- for given electron temperature T_e , velocity match only for ion energies

$$K_i \lesssim \frac{m_i}{m_e} k T_e$$

* $T_e = 4 \text{ K} \rightarrow K_i \lesssim 1 \text{ eV/q}$ * $T_e = 300 \text{ K} \rightarrow K_i \lesssim 100 \text{ eV/q}$

• at higher energies, electron cooling is 'safe' from recombination

• at lower ion energies, recombination is essentially a function of the electron temperature

$$\Gamma_{\rm RR} \approx 6 \times 10^{-11} \frac{n_e}{T^{1/2}} \mathrm{s}^{-1} \times Z^2$$

- with $n_e = 10^7$ cm⁻³, $T_e = 300$ K, and Z = 50: $\Gamma_{\rm RR} \approx 0.09~{\rm s}^{-1}$
- this looks feasible !

G. Werth, 2002 GSI HITRAP workshop

Electron Cooling in Penning Trap: Surviving Probability of U⁹²⁺ Ions

T = 4 K



Proton cooling of HCI

 similar to electron cooling, however no self-cooling of protons (or He²⁺) → need initially cold sample

• Simulations by P. Grothkopp



Fig. 2 Uranium (90+) with an initial energy of 200 eV/q and different numbers of injecting ions from 10 to 10⁴. Cooled by protons with a density of $n = 10^{7}$ cm⁻³ and a number of $N = 10^{7}$.

At the Brillouin limit, cooling could be really fast!



Fig. 5 Uranium (90+) with an initial energy of 50 eV/q and different numbers of injecting ions from 10 to 10⁴. Cooled by protons with a density of $n = 4.23 \times 10^{10} \text{ cm}^{-3}$ and a number of $N = 10^{-7}$ cool ions.

• no self-cooling $\Rightarrow T_{eq} \approx T_{coolant}$ only for $N_{HCI} \ll N_p$

Design of a HCI cooler trap

• what has been done in this field so far ?

- nested Penning trap for (anti)proton cooling with electrons / positrons (Gabrielse et al., also ATHENA)
- cooler Penning trap for HITRAP project at GSI
- RIKEN HCI cooler trap for surface studies

The HITRAP cooler trap at GSI (under development)



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continue with resistive cooling to go to 4K





(b) e⁺ Accumulation



(c) e⁺/e⁻ Cooling of HCIs



(d) Extraction of Cold HCIs





Propose to build a cooler trap capable of electron cooling and light-ion cooling

- design principles can be very similar to HITRAP and RIKEN cooler traps
 - long trap structure (≈ 50 cm)
 - * accommodate ion bunches from EBIT (25 cm long)
 - * form multiple nested potentials
 - \star large diameter ($\gtrsim 1$ ")
 - current design: shoot through cooler trap
 - * off-axis insertion of electrons/protons (RIKEN trap)

6 T magnet for electron cooling 10 ppm magnet, 4-5" warm bore 1 ppm or better for addt'l applications ? TITAN Penning trap at room temperature make system cryo-ready, resistive cooling etc. to 4K

• integrate 'mini-EBIT' for offline tests of cooling ?

... finally the sad part

update, June 20, 2005: finally the good news the CFI proposal has been approved by the province !