



**TITAN**  
ISAC-TRIUMF

S1283

**Precision Mass  
Measurements of the Halo  
Candidates  $^{31}\text{Ne}$  and  $^{22}\text{C}$**

stephan ettenauer  
for the TITAN collaboration



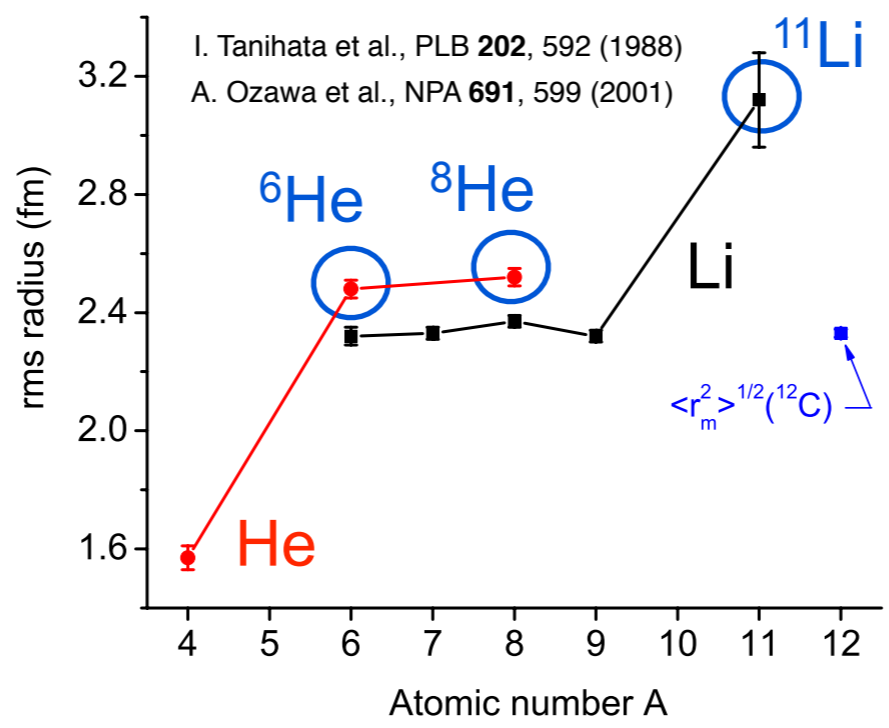
SAP-EEC, July 2010

# Halo Nuclei

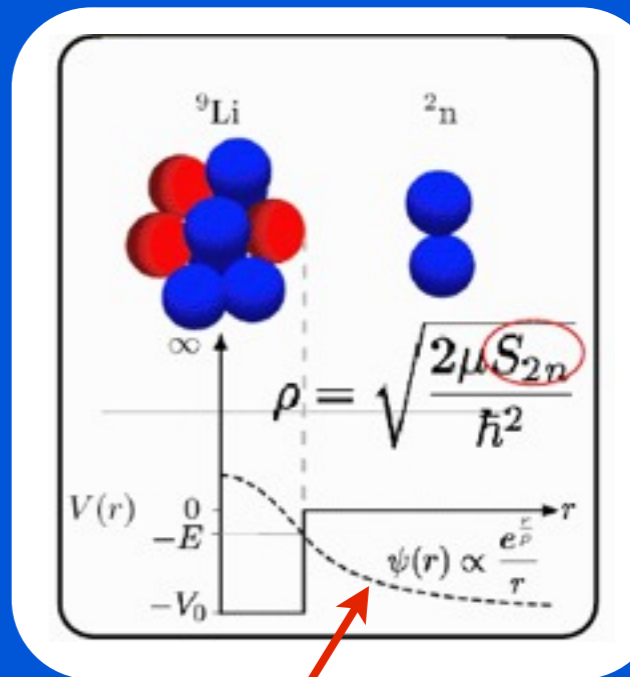
extreme n/p ratios

Halo	n/p
<sup>6</sup> He	2
<sup>8</sup> He	3
<sup>11</sup> Li	2.66
<sup>14</sup> Be	2.5
<sup>19</sup> C	2.17
<sup>12</sup> C	1

large radii

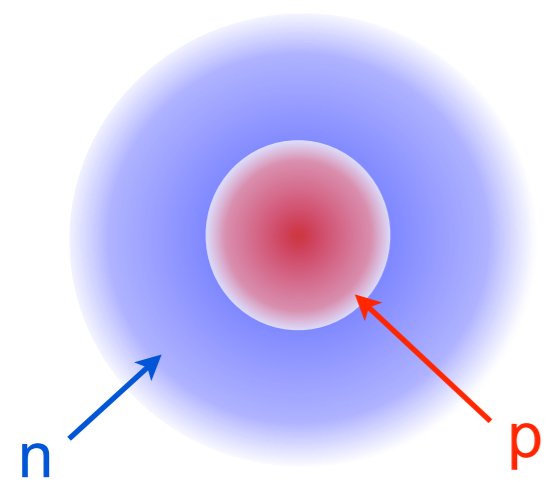


nucleons in classically forbidden region

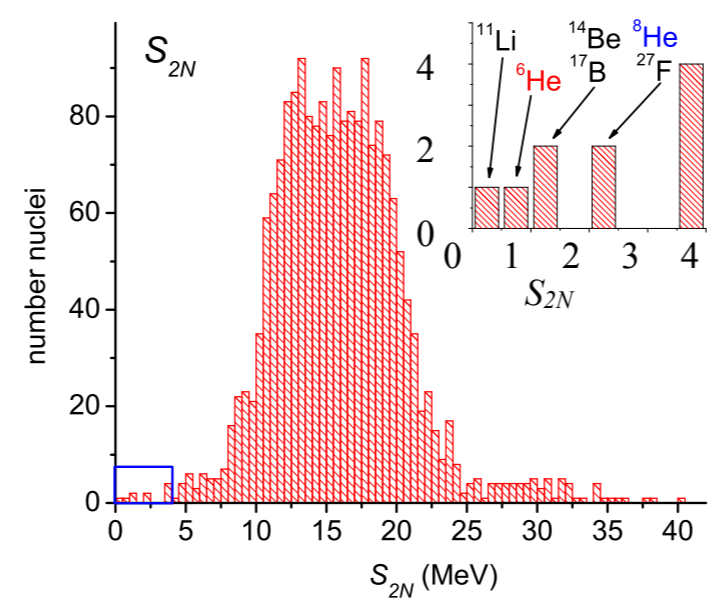


exp. fall off

but  $R_{\text{matter}} \neq R_{\text{charge}}$



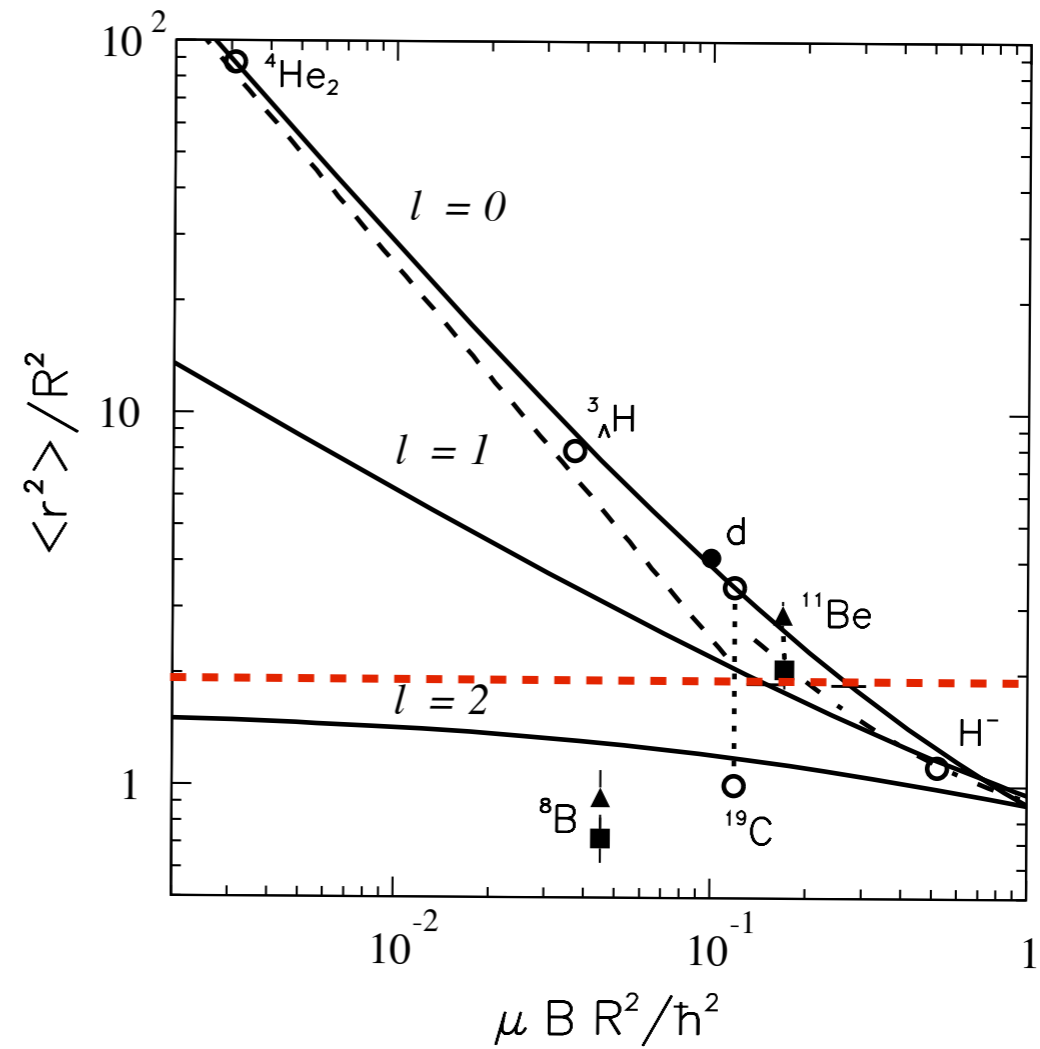
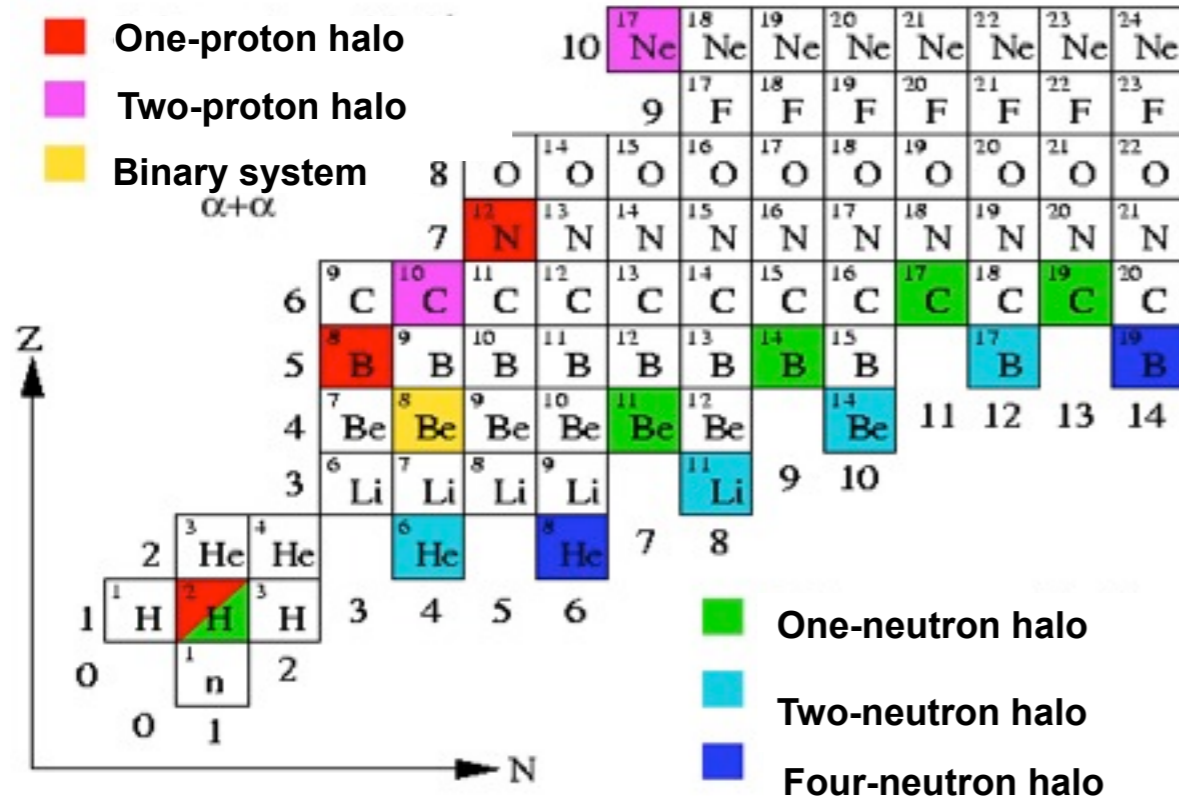
tiny separation energies



due to exotic features:

ideal test ground for nuclear structure theory

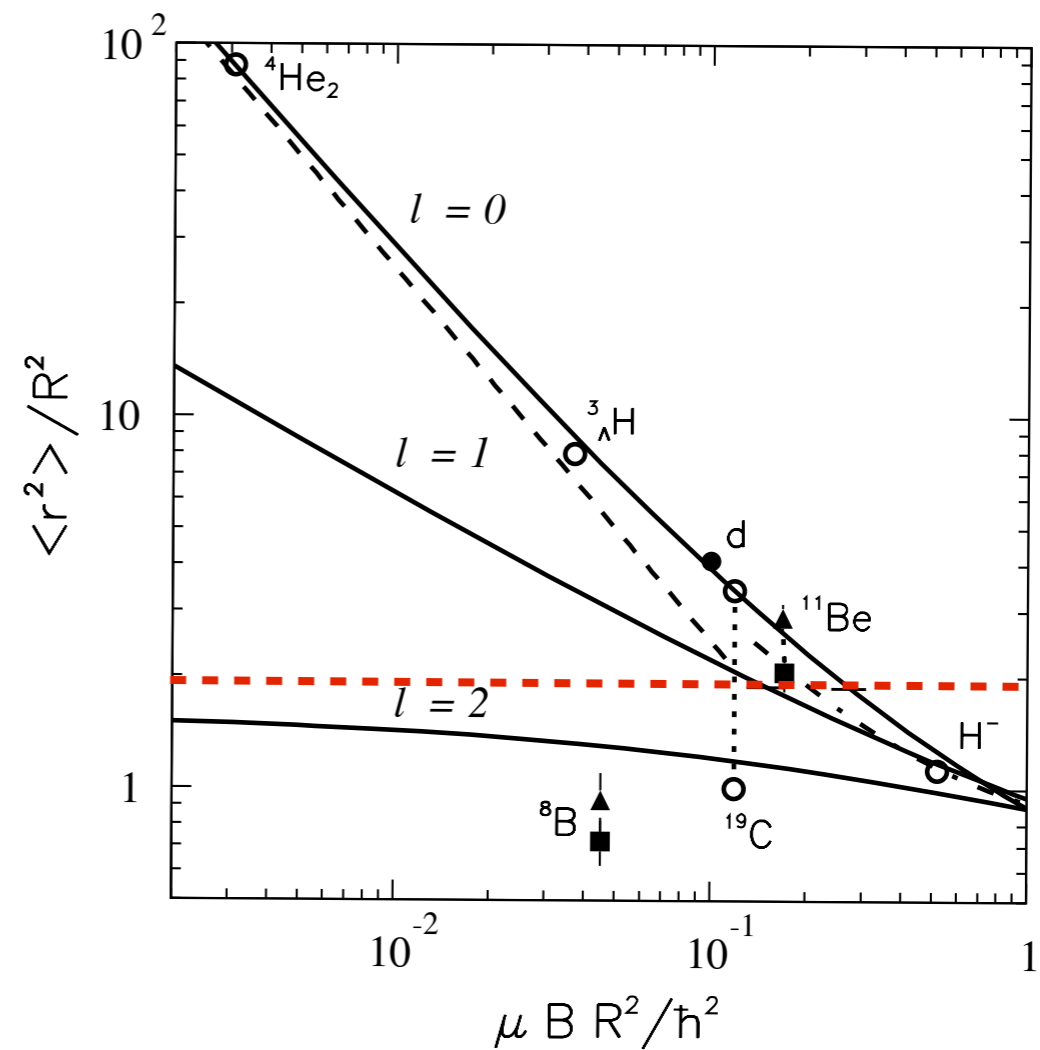
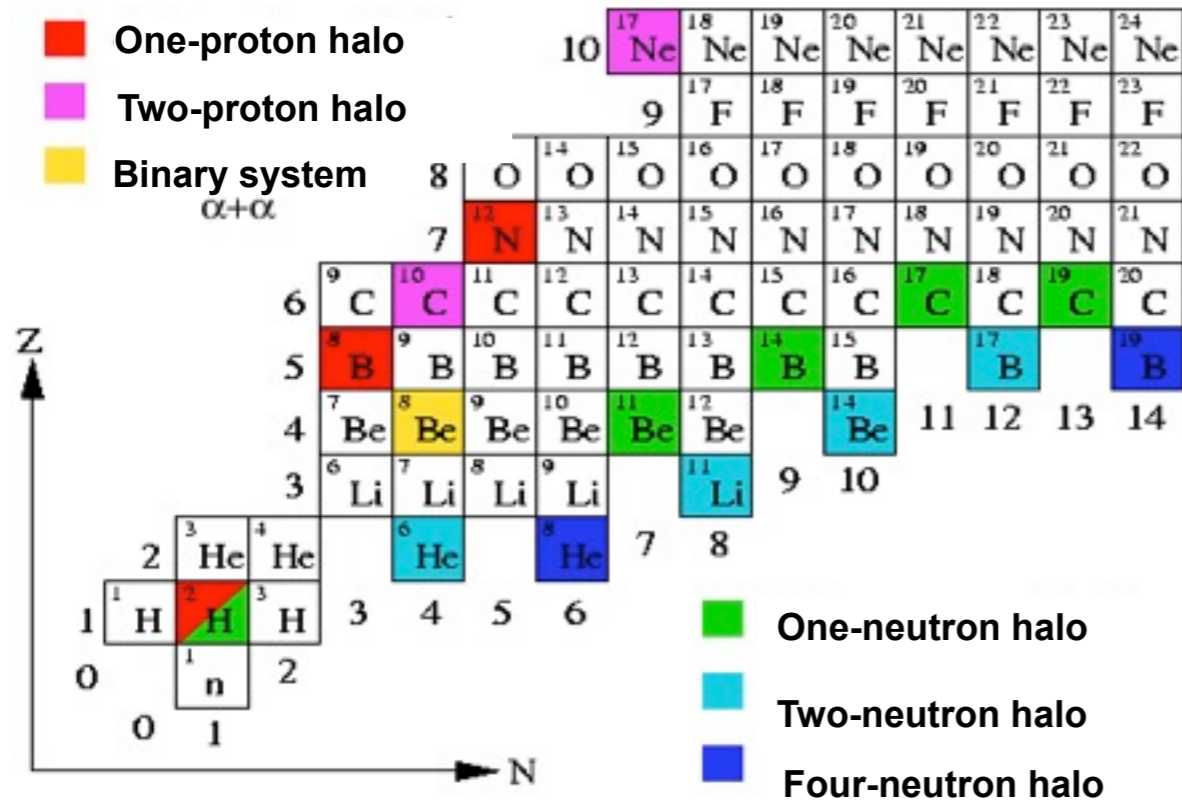
# heavier halos?



A. S. Jensen et al., Rev. Mod. Phys. 76, 215–261 (2004)

- ⇒ halos only for low orbital angular momentum (s and p)
- ⇒ generally not along the drip-line in heavier systems
- ⇒ halos limited to lighter masses ?

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A. S. Jensen et al., Rev. Mod. Phys. 76, 215–261 (2004)

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## potential mechanisms for low $\ell$ at heavier A:

- shell evolution
- deformation & halos
- halo formation adds extra stability ?

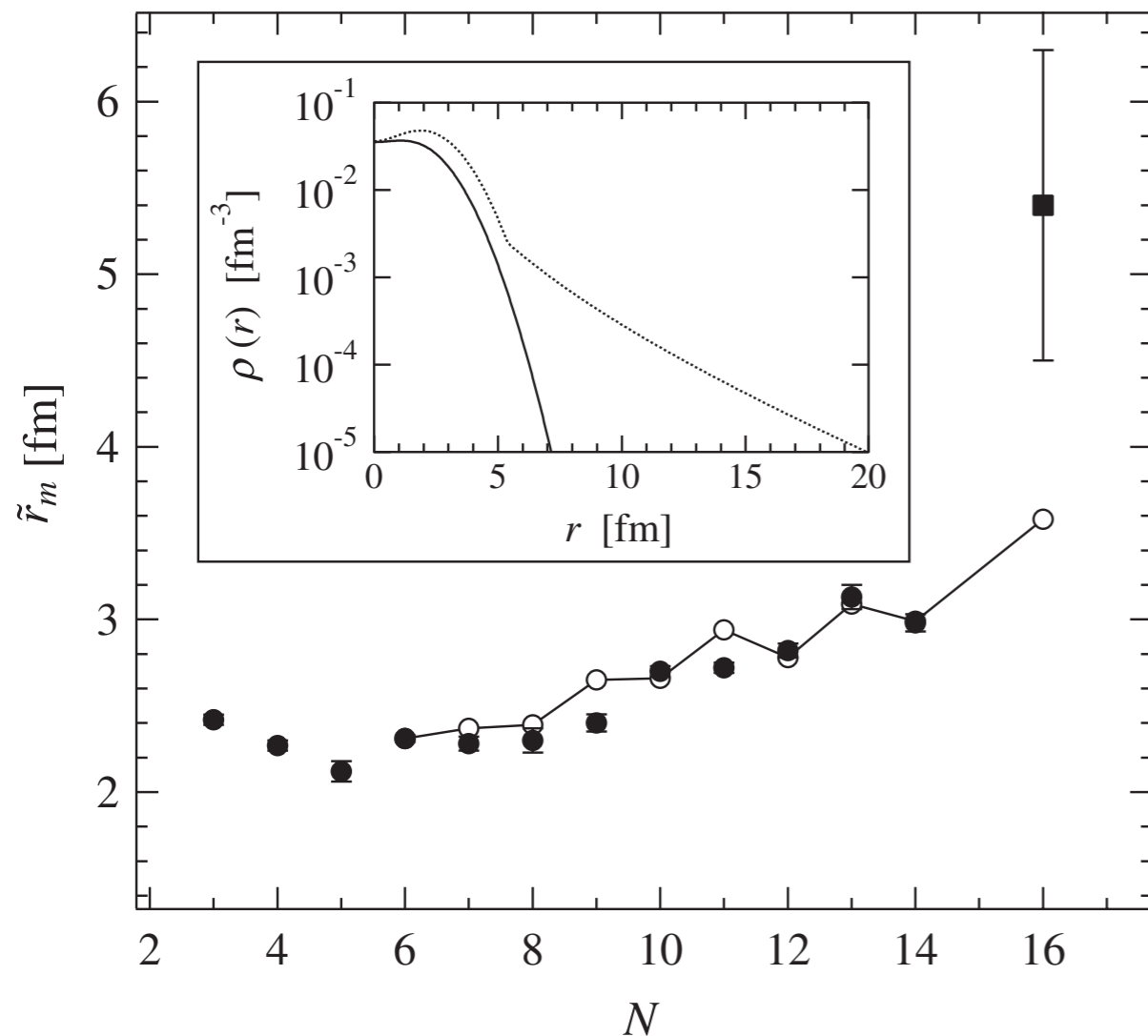
halos more common than expected?

# reaction cross section of $^{22}\text{C}$

- transmission experiment @ RIKEN
- $^{19,20,22}\text{C}$  on liquid H-target
- 40A MeV

TABLE I. Reaction cross sections ( $\sigma_R$ ) in millibarns.

$A$	$\sigma_R$	$\sigma_R^{\text{calc.}}$ [22]
19	754(22)	758
20	791(34)	761
22	1338(274)	$\geq 957$

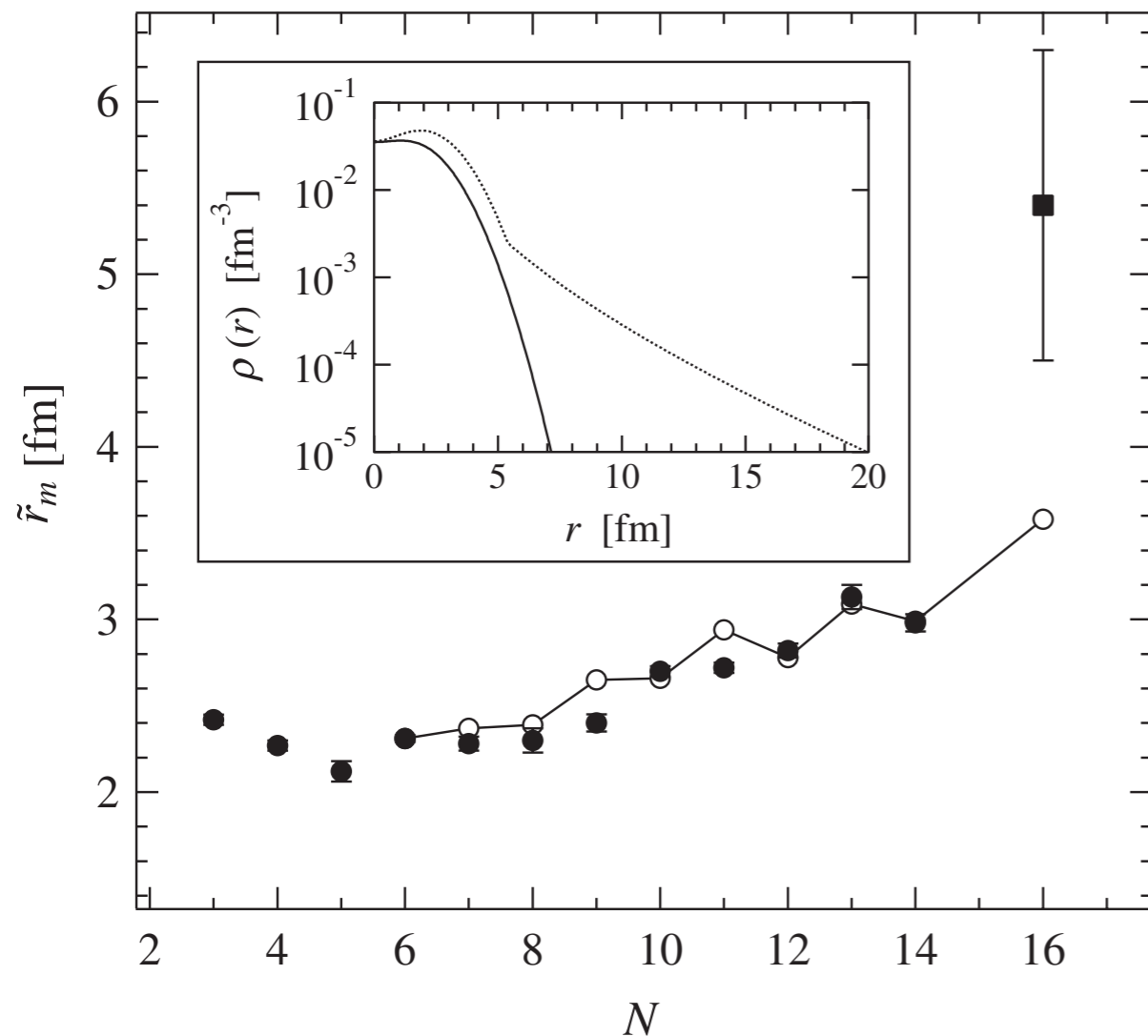


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## Glauber type calculation:

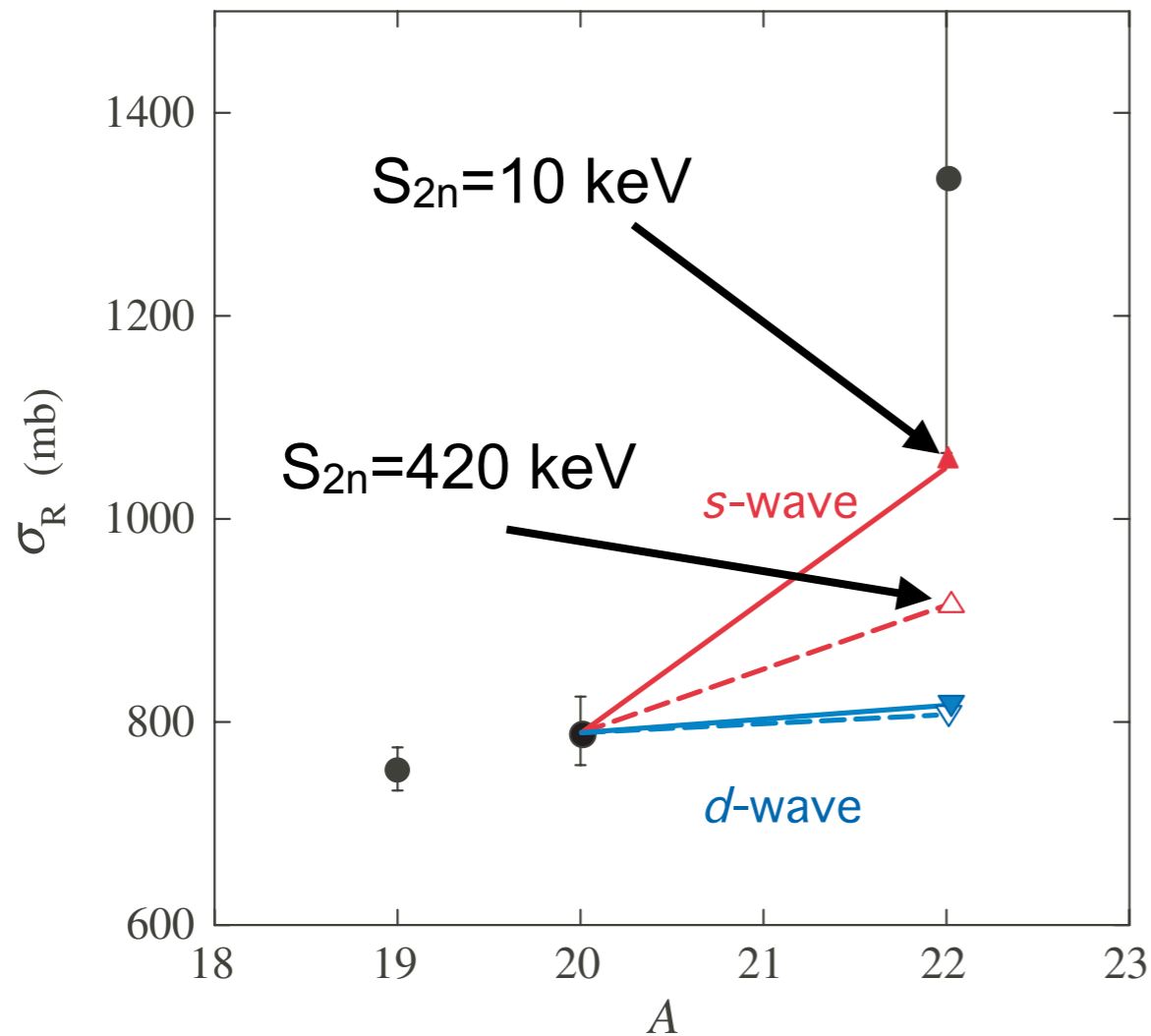
- optical limit ( $^{20}\text{C}$ )
- FB approach
  - ▶  $^{18}\text{C}+n$
  - ▶  $^{20}\text{C}+n+n$  (s-wave)

$S_{2n}$ (MeV)	$r_m$ (fm)	$\sigma_R$ (mb)
0.489	3.6	957
0.361	3.7	969
0.232	3.8	985
0.122	4.1	1005

# reaction cross section of $^{22}\text{C}$

## Glauber calculation for $^{22}\text{C}$

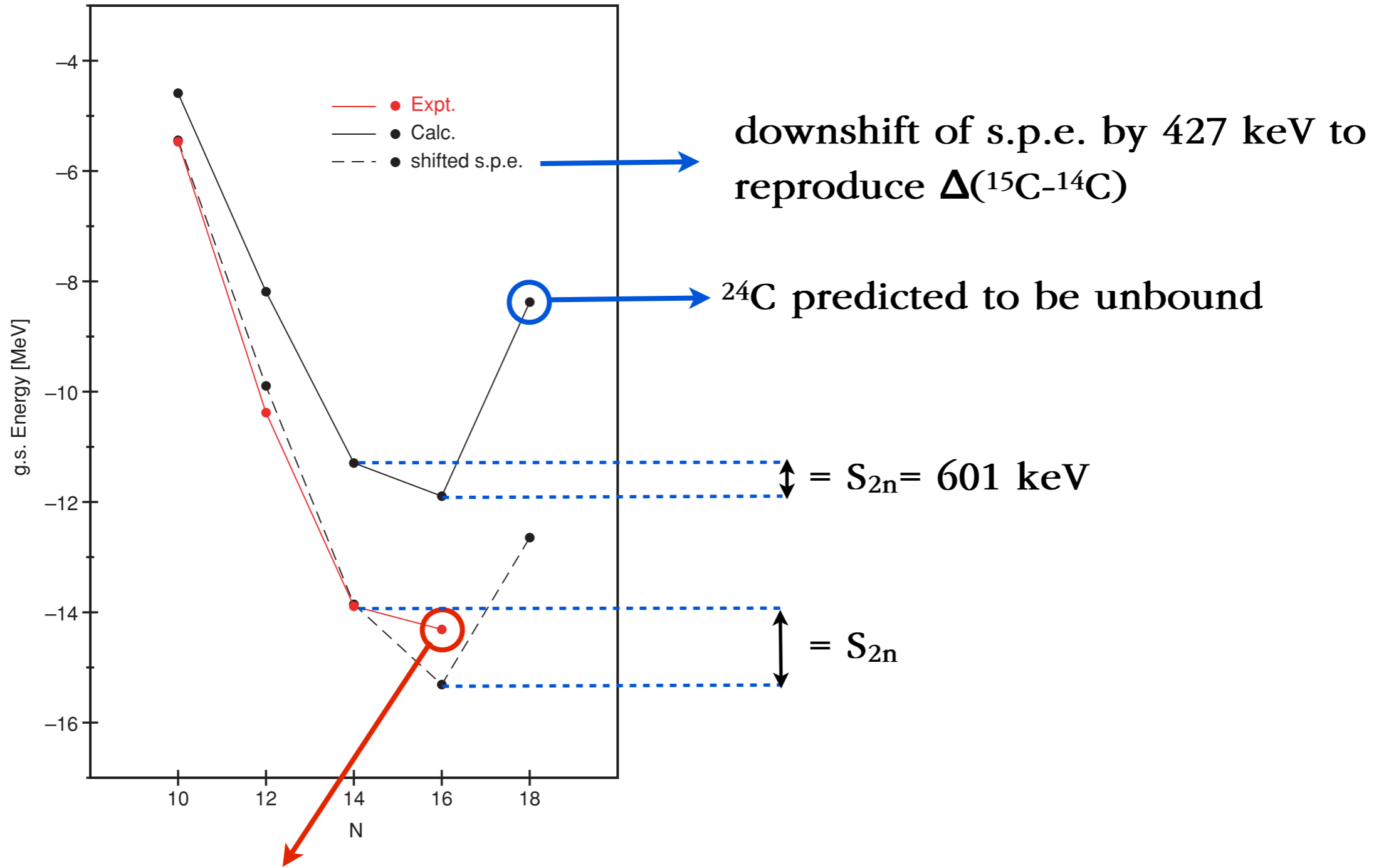
- FB approach
- finite-range treatment
- vary  $S_{2n}$
- spectroscopic factor  $(0d5/2)^2_{J=0} \leftrightarrow (1s_{1/2})^2_{J=0}$



$$S_{2n} = 420 \pm 920 \text{ keV}$$

- currently extrapolation in AME'03
- reaction data favours lower  $S_{2n}$
- larger  $S_{2n}$ :  
**validity of Glauber model at 40A MeV ?**

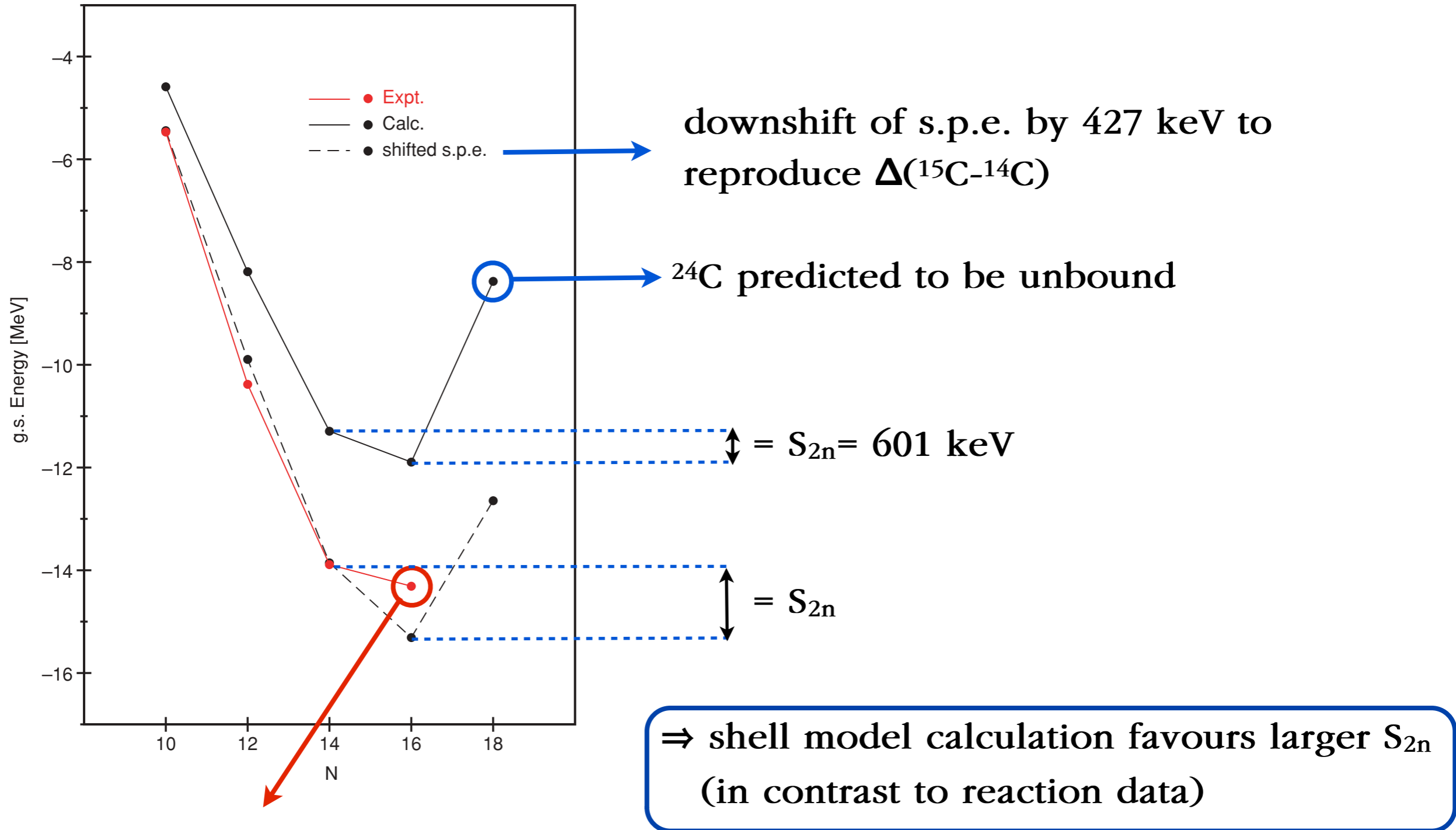
# shell model calculation for $^{22}\text{C}$



AME'03 prediction



# shell model calculation for $^{22}\text{C}$



AME'03 prediction

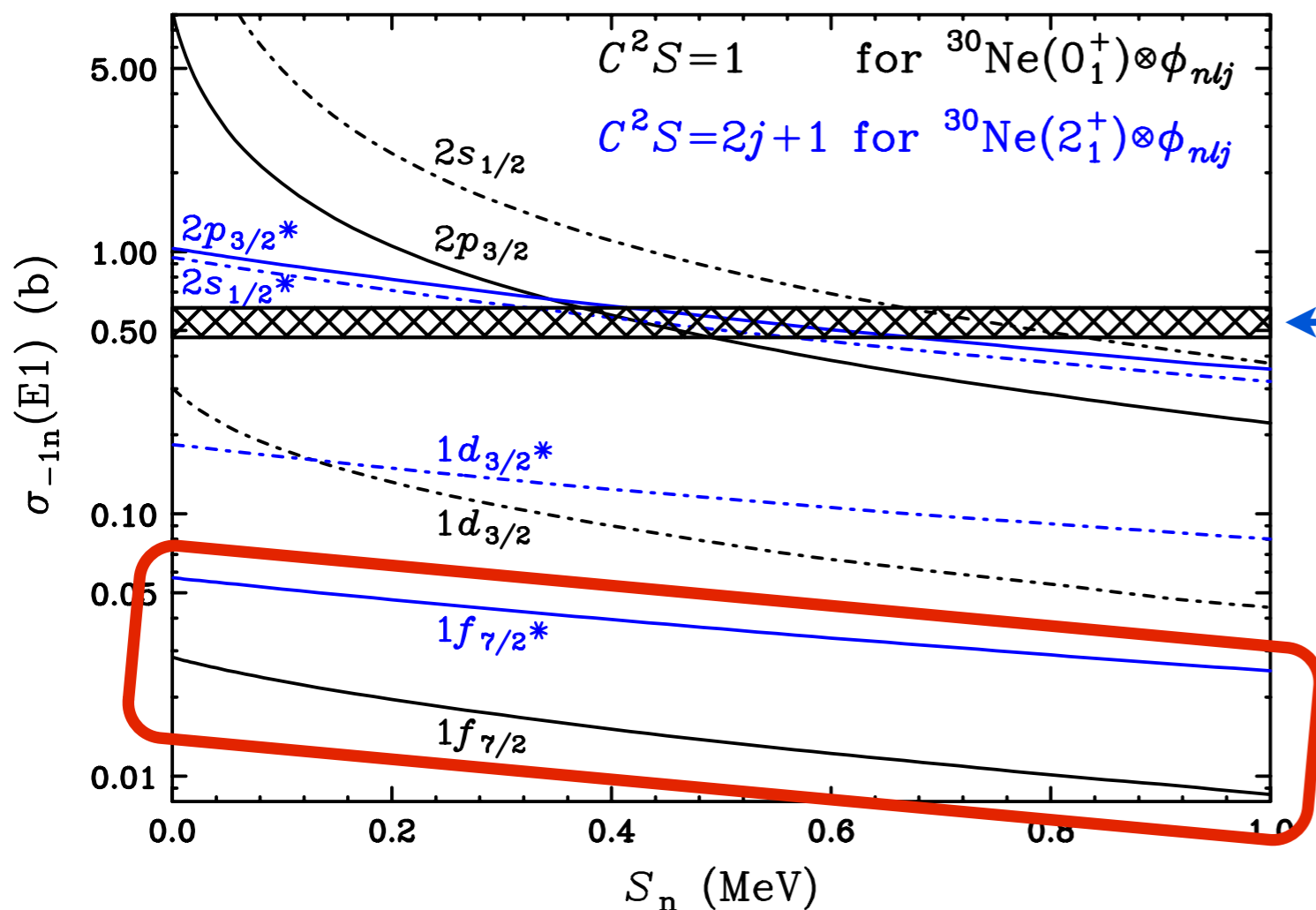
# Coulomb breakup of $^{31}\text{Ne}$ (RIKEN)

TABLE I. Single-neutron removal cross sections ( $\sigma_{-1n}$ ) for  $^{31}\text{Ne}$  and  $^{19}\text{C}$  on Pb and C targets at the incident energies shown. The ratio of the measured cross sections and the deduced Coulomb breakup cross sections are also listed.

Reaction	$\bar{E}/A$ (MeV)	$\sigma_{-1n}$ (mb)	$\frac{\sigma_{-1n}(\text{Pb})}{\sigma_{-1n}(\text{C})}$	$\sigma_{-1n}(E1)$ (mb)
$^{31}\text{Ne} + \text{Pb}$	234	712(65)	9.0(1.1)	540(70)
$^{31}\text{Ne} + \text{C}$	230	79(7)		
$^{19}\text{C} + \text{Pb}$	243	969(34)	7.4(4)	690(70)
$^{19}\text{C} + \text{C}$	238	132(4)		

$$S_n = 0.3 \pm 1.6 \text{ MeV}$$

*B. Jurado et al., Phys. Lett. B 649, 43 (2007)*

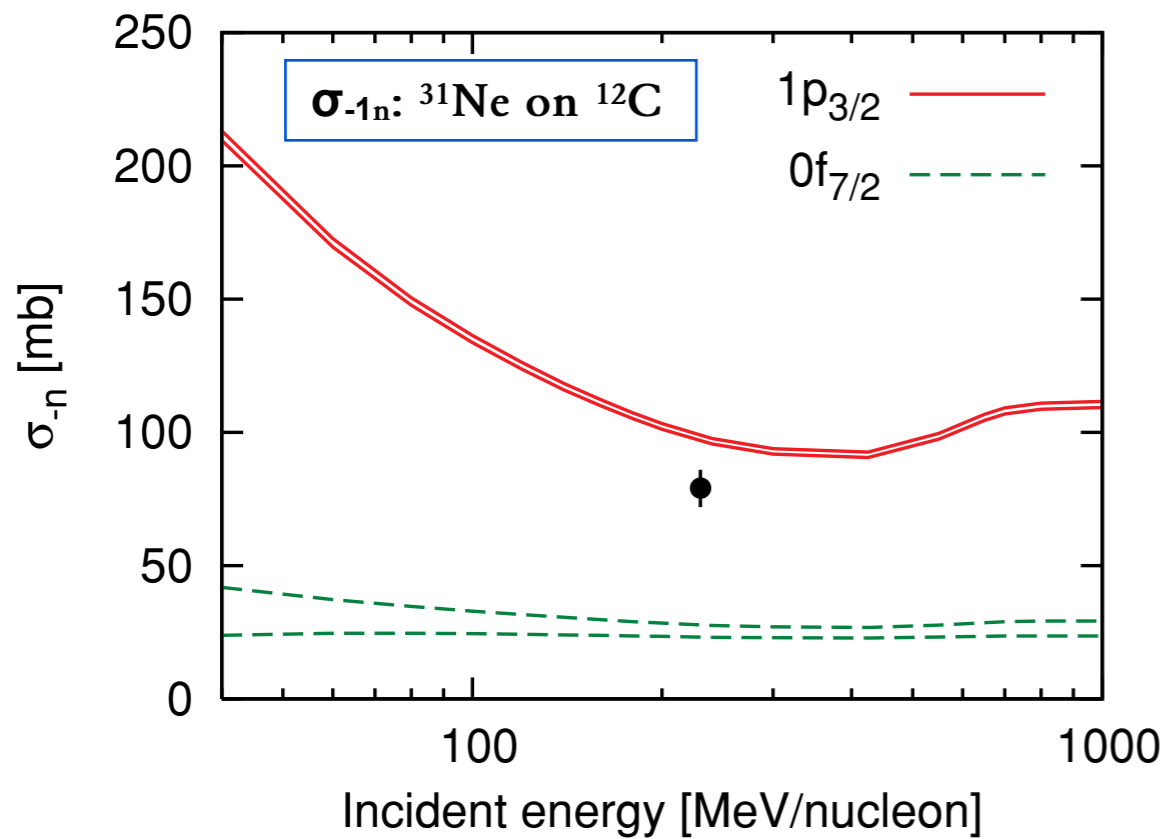


experiment:  
consistent with formation of halo

expected from shell-model

# Glauber calculation $\sigma_{-1n}$

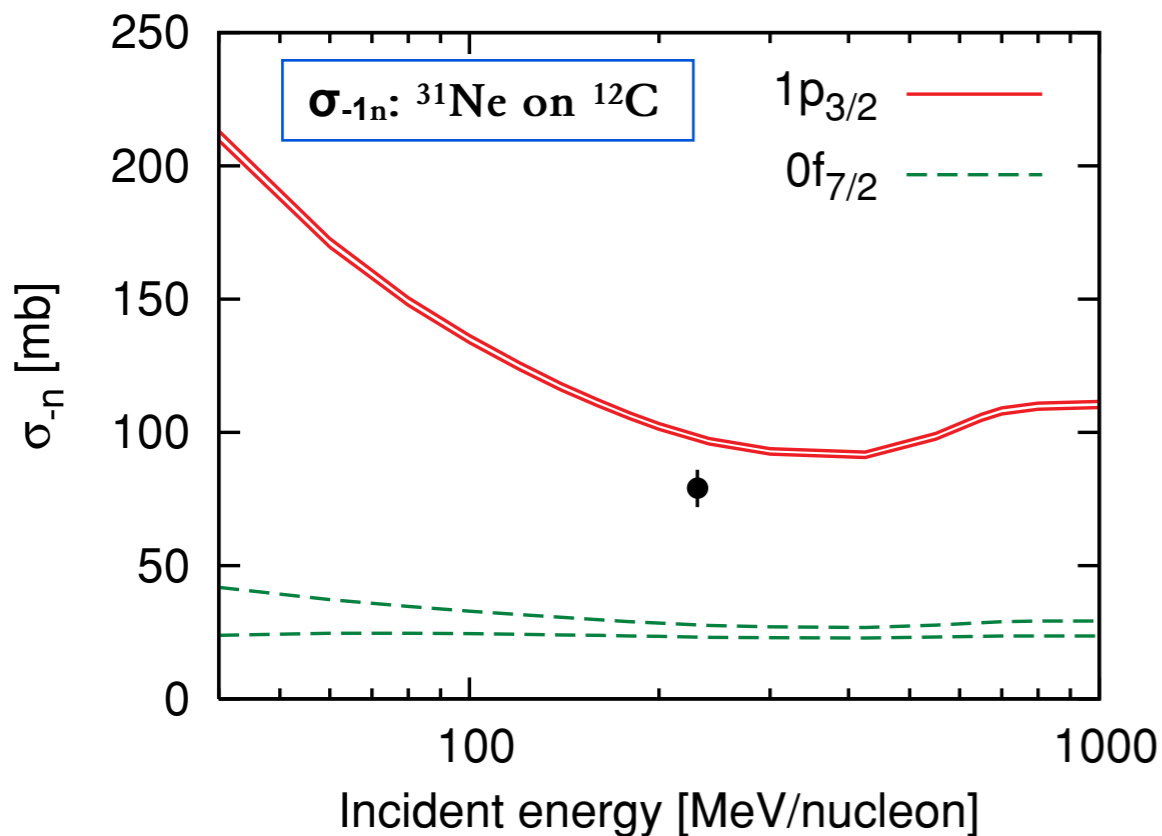
$${}^{30}\text{Ne}(0^+) \otimes 1p_{3/2} \leftrightarrow {}^{30}\text{Ne}(0^+) \otimes 0f_{7/2}$$



[mb]	exp.	th.: $p_{3/2}$	th.: $f_{7/2}$
$\sigma_{-1n}$ on ${}^{12}\text{C}$	79(7)	96	26
$\sigma_{-1n}$ on ${}^{208}\text{Pb}$	712(65)	1140	91

# Glauber calculation $\sigma_{-1n}$

$${}^{30}\text{Ne}(0^+) \otimes 1p_{3/2} \leftrightarrow {}^{30}\text{Ne}(0^+) \otimes 0f_{7/2}$$



n removal cross section at ~240A MeV:

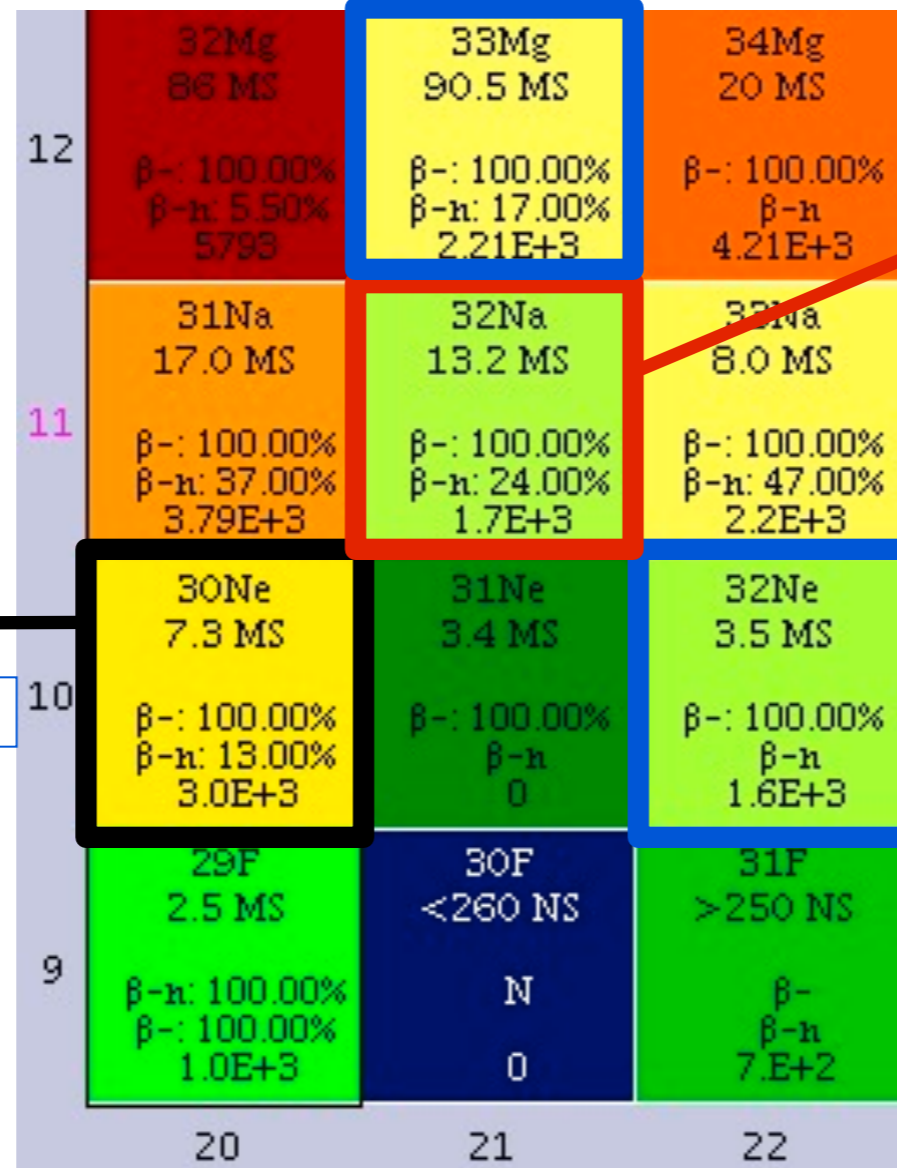
[mb]	exp.	th.: $p_{3/2}$	th.: $f_{7/2}$	$S_n$ [keV]
$\sigma_{-1n}$ on ${}^{12}\text{C}$	79(7)	96	26	
$\sigma_{-1n}$ on ${}^{208}\text{Pb}$	712(65)	1140	91	0.33
		750		0.6

⇒ strong dependence on  $S_n$

# Island of Inversion

consistent with  $I^\pi=3/2^-$

*D. T. Yordanov et al., PRL 99,212501 (2007)*



mass measurement

*C. Thibault et al., Phys. Rev. C 12, 644 (1975)*

low  $2^+$  state

*Y. Yanagisawa et al., PLB 566, 84 (2003)*

low  $2^+$  state

*P. Doornenbal et al., PRL 103, 032501 (2009)*

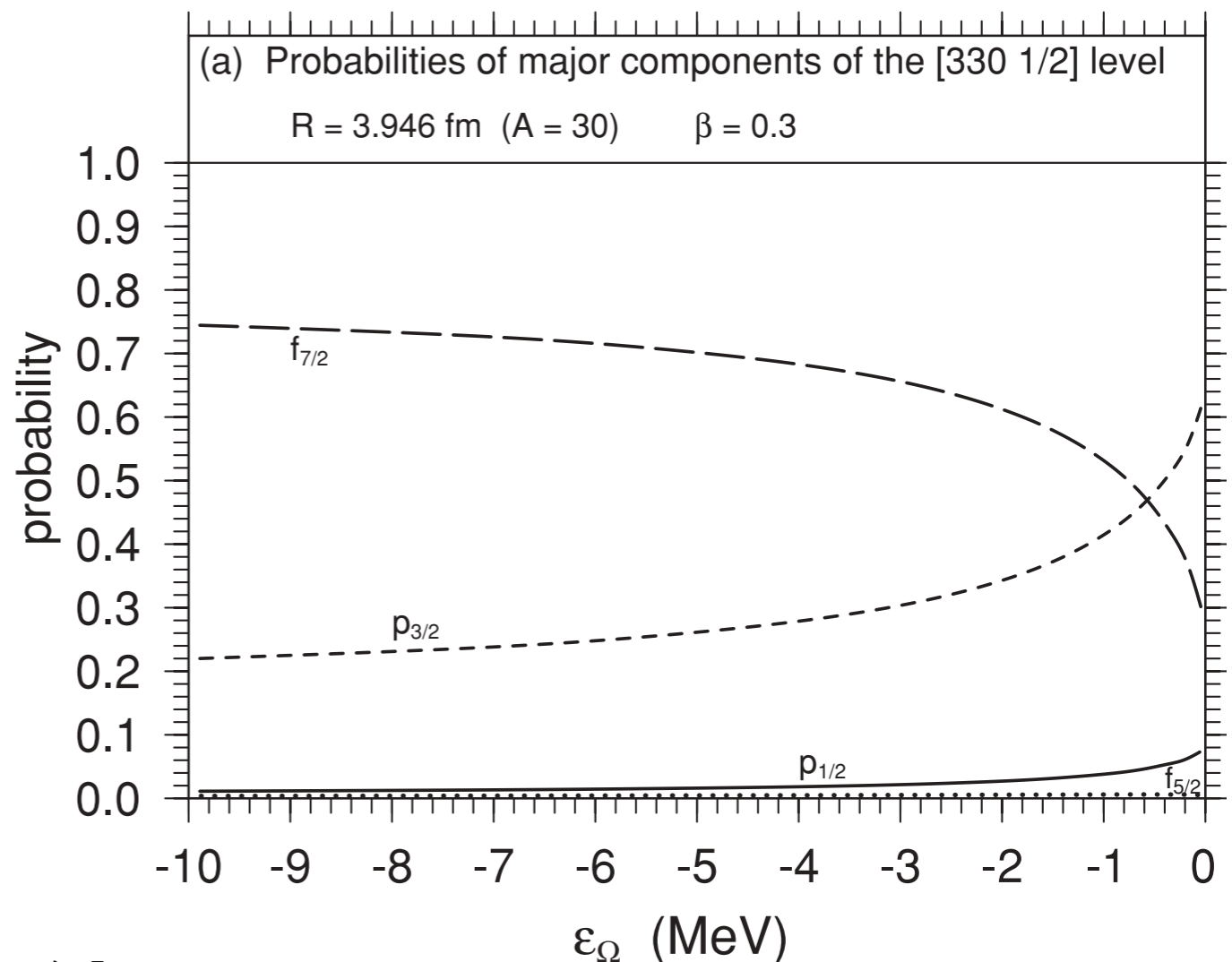
$\Rightarrow$   $^{31}\text{Ne}$  likely to be deformed

# Deformation in $^{31}\text{Ne}$

- axially symmetric quad-deformed Wood-Saxon
- dominance of lowest  $\ell$ -component for low binding energy
- irrespective of size of deformation or kind of orbit

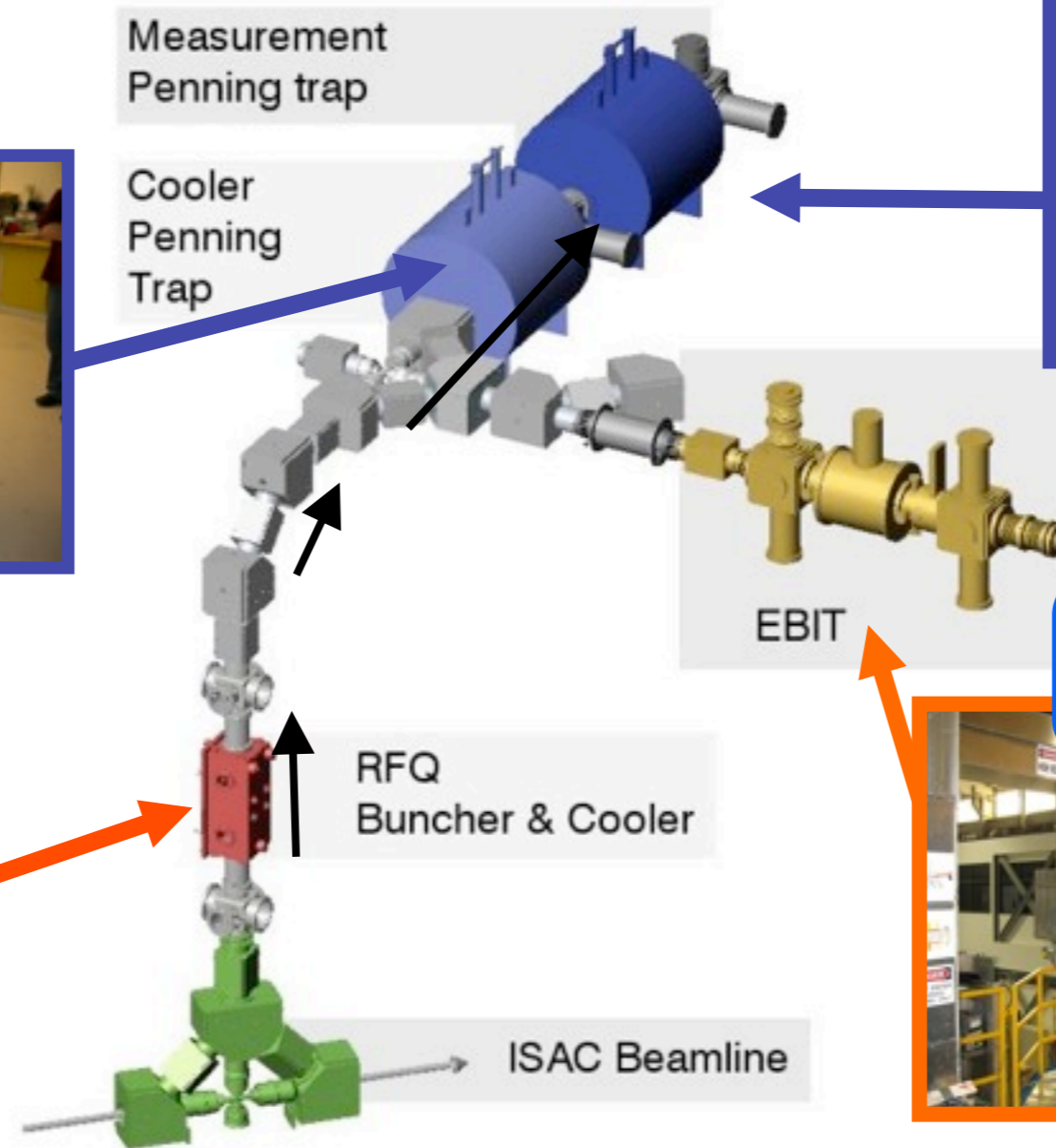
## Conclusions for $^{31}\text{Ne}$ :

- $S_n > 500$  keV
  - the deformation large ( $\beta > 0.6$ )
  - $I^\pi = 1/2^+$  from the Nilson level [200 1/2].
- $S_n < 500$  keV
  - $I^\pi = 3/2^-$  and a p-wave neutron halo coming from either
  - [330 1/2] for  $200 \text{ keV} < S_n < 500 \text{ keV}$  or
  - [321 3/2] for  $200 \text{ keV} > S_n$

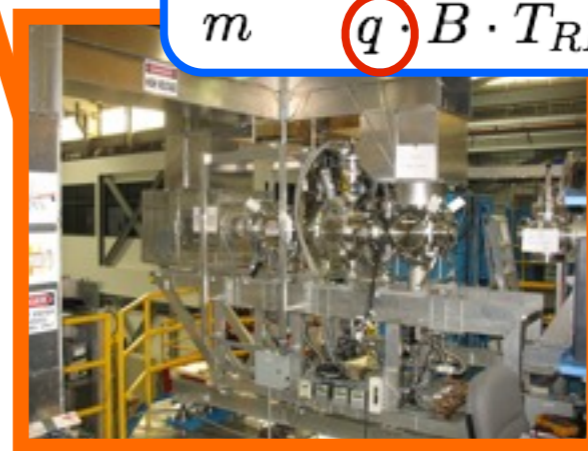


determination of separation energies via direct mass measurement

$$S_n = m(Z, N - 1) + m_n - m(Z, N)$$



$$\frac{\delta m}{m} \approx \frac{m}{q \cdot B \cdot T_{RF} \cdot \sqrt{N_{ion}}}$$



ISAC beam: A<sup>+</sup> →

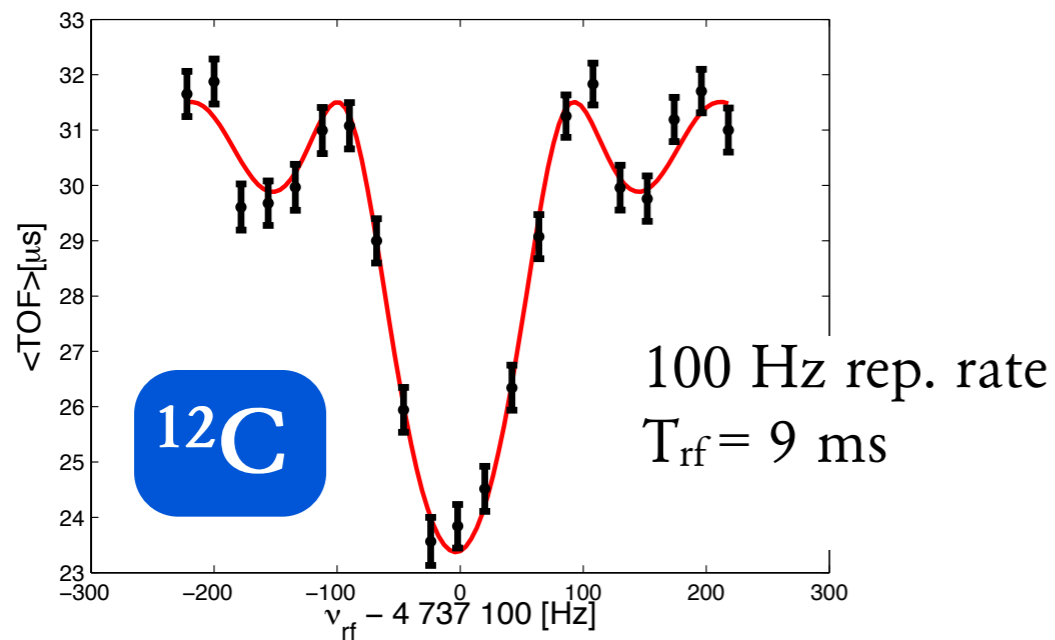
# Measurement @ TITAN

## short half-lives:

- 3.4(8) ms for  $^{31}\text{Ne}$  and 6.1(1.4) ms for  $^{22}\text{C}$
- $^{11}\text{Li}$ : 8.8 ms

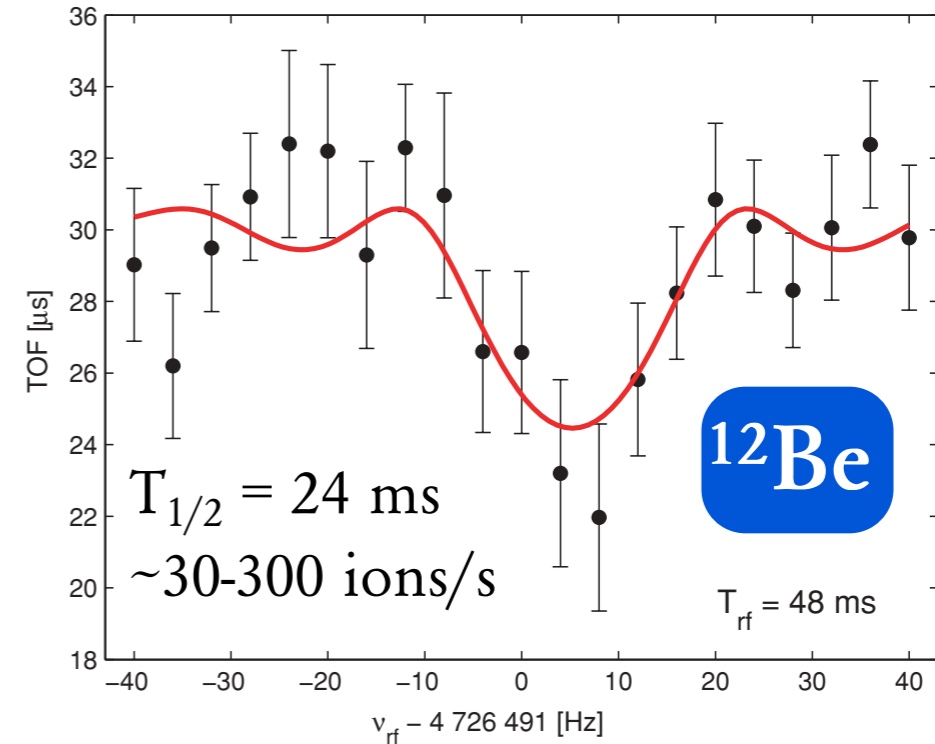
*M. Smith et al., PRL 101, 202501 (2008)*

- $T_{\text{rf}} \approx 2.9 \cdot T_{1/2} \Rightarrow T_{\text{rf}} = 10 \text{ ms}$



## low yields:

m.e.=25 078.0(2.1) keV



*S. Ettenauer et al., PRC 81, 024314 (2010)*

**propose:** mass measurement of  $^{20,22}\text{C}$  and  $^{31,30}\text{Ne}$  ( $^{30}\text{Ne}$  part of S1240) with  $\delta m/m \approx 5 \cdot 10^{-7}$

**request:** 3 shifts + 1 shift setup for each case (12 shifts total)

requires beam development (UO or UC for Ne, UO for C)  $\Rightarrow$  stage 1

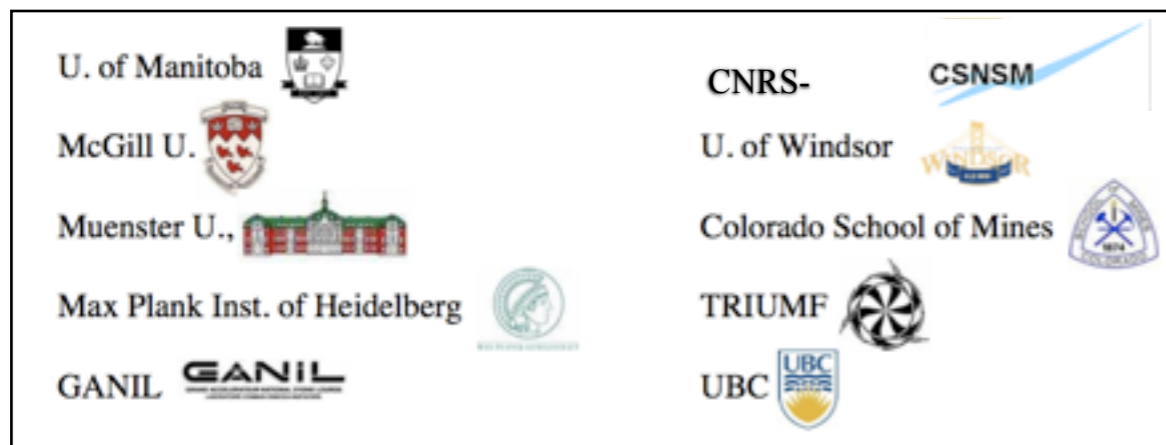
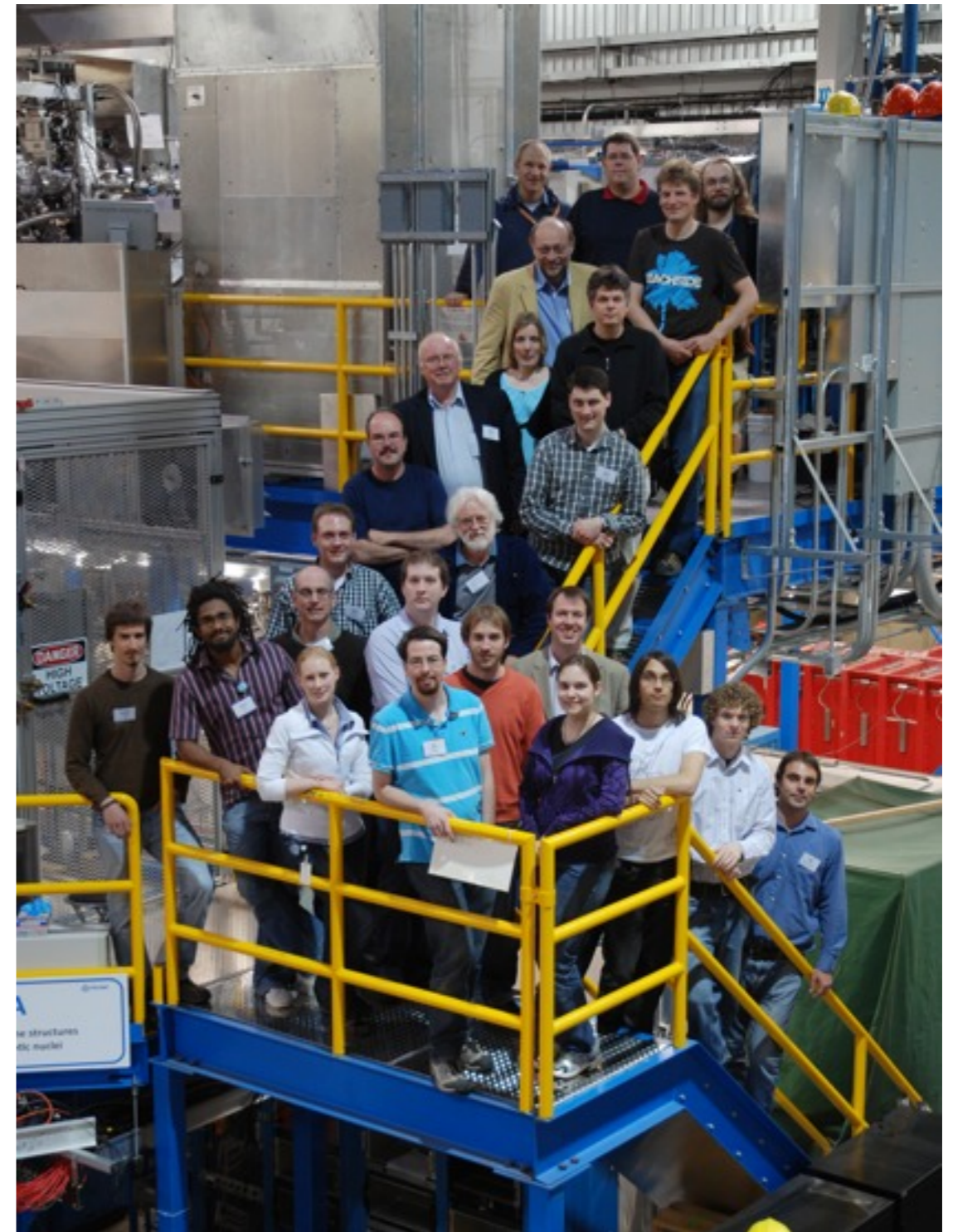


# Summary

- halo nuclei
  - ideal testing grounds for nuclear structure models
  - limited to lighter masses?
  - mechanism for halos in heavier systems?
- recent reaction measurements
  - $^{22}\text{C}$  in transmission experiment
  - $^{31}\text{N}$  in Coulomb breakup
 } evidence for halo structures in heavier systems
- in both cases: knowledge on  $S_{2n}$  limited
  - ⇒ interpretation of data ambiguous
  - ⇒ uncertainty in reaction models
- for  $^{22}\text{C}$ : mass also benchmark for shell-model calculation
- TITAN:
  - fastest Penning trap spectrometer:  $T_{1/2} < 10$  ms possible
  - measurements with low yields feasible
  - propose mass measurement of  $^{20,22}\text{C}$  and  $^{31,30}\text{Ne}$  to extract  $S_{2n}$  and  $S_n$
  - request 12 shifts / stage 1

# S1283 collaboration

- ❖ TRIUMF: J. Dilling, P. Delheij, M. Pearson, E. Mané, M. C. Simon, T. Brunner, U. Chowdhury, B. Eberhart, S. Ettenauer, A. Gallant, V. Simon
- ❖ CSNSM-IN2P3-CNRS: D. Lunney
- ❖ University of Manitoba: G. Gwinner
- ❖ NSCL: M. Brodeur, R. Ringle



# Backup Slides

# Precise & Accurate

line width (FWHM):

$$\Delta\nu \approx 1/T_{rf}$$

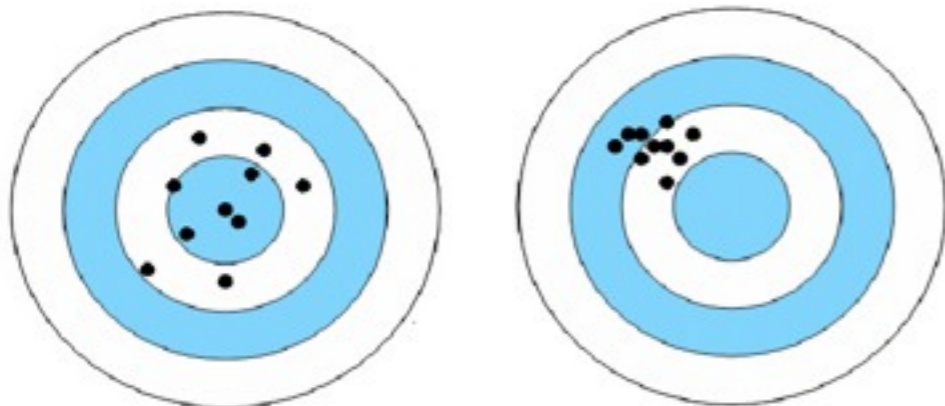
⇒ resolution:

$$R = \frac{m}{\Delta m} = \frac{\nu_c}{\Delta\nu_c} \approx \nu_c T_{rf}$$

$$\approx \frac{qBT_{rf}}{2\pi m}$$

⇒ even for  $T_{rf} \sim 10\text{ms}$

$$(\delta m/m)_{\text{stat}} < 10^{-7}$$



accurate,  
but not precise

precise,  
but not accurate

- exact theoretical description

*L.S. Brown and G. Gabrielse, Rev. Mod. Phys. 58, 233 (1986)*  
*G. Bollen et al., J. Appl. Phys. 88, 4355 (1990)*  
*M. König et al., Int. J. Mass Spect. 142, 95 (1995)*  
*M. Kretschmarr, Int. J. Mass Spect. 246, 122 (2007)*

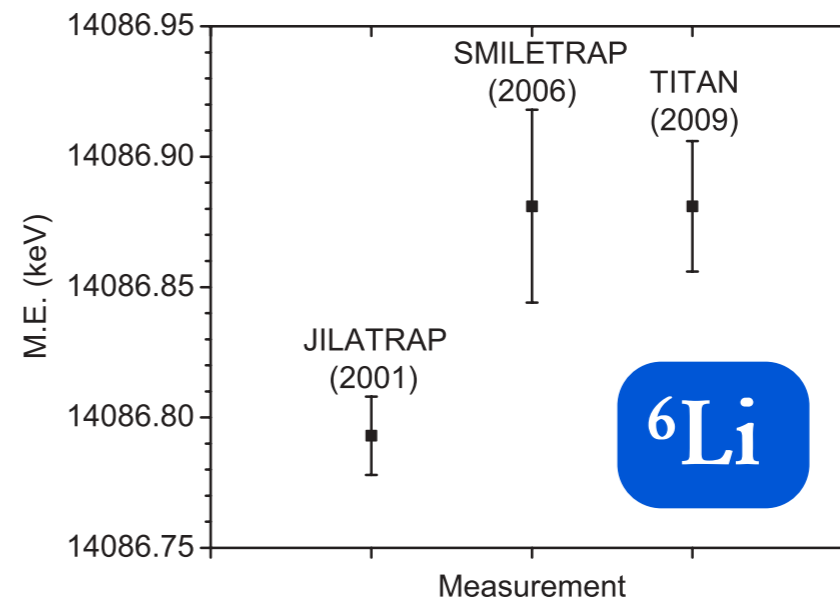
- even for non-ideal traps

*G. Bollen et al., J. Appl. Phys. 88, 4355 (1990)*  
*G. Gabrielse, PRL 102, 172501 (2009)*

- off-line tests with stables

⇒ control over systematics

for TITAN: < 5 ppb possible



*M. Brodeur et al, PRC 80, 044318 (2009)*

# Coulomb Breakup

- low-lying electric dipole strength E1 is enhanced for halos, because:
- one neutron-removal cross section  $\sigma_{-1n}(E1)$  due to dipole strength E1 follows

$$\sigma_{-1n}(E1) = \int_{S_n}^{\infty} \frac{16\pi^3}{9\hbar c} N_{E1}(E_x) \frac{dB(E1)}{dE_x} dE_x$$

## E1 virtual photon number

- exponentially decreasing

## E1 strength distribution

- peaked at around 1 MeV for s and p
- maximum at higher  $E_x$  otherwise

- ▶ enhanced for s and p orbitals
- ▶ estimating  $\sigma_{-1n}(E1)$  is thus a signature for a halo

- measured  $\sigma_{-1n}$  has contributions from Coulomb and from nuclear interaction
- 2 targets:

- Pb (Z=82): Coulomb dominated
- C: pure nuclear contribution
- $\Gamma$  scales nuclear part from C to Pb

$$\sigma_{-1n}(E1) = \sigma_{-1n}(Pb) - \Gamma \sigma_{-1n}(C)$$

